

## Today (Feb 4):

- complete: strain and rotation from displacement gradient matrix
- examples
- GPS

## Group activity - This will be part 1 of your Homework 3 (hold onto it)

Trace axes and points onto the transparency.

Rotate the transparency (not by a large amount).

1) For the point on the  $x_1$  axis: what is  $du_2$ ? What is  $dx_1$ ? ( $dx_2$  and  $du_1$  should be small enough to ignore.)

2) For the point on the  $x_2$  axis: what is  $du_1$ ? What is  $dx_2$ ? (careful with signs)

Remember the definition of tangent of  $\theta$  and that for a small angle,  $\theta$  in radians =  $\tan(\theta)$

3a) What is the angle of rotation ( $\theta$ ) of the  $x_1$  axis? Of the  $x_2$  axis?

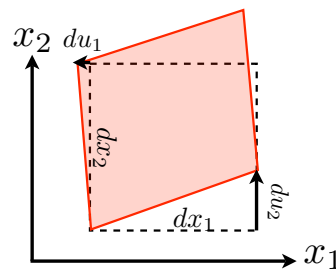
3b) We want to define counter-clockwise rotation as positive, then what are the answers to 3a?

3c) Write out the mean of the two angles in 3b (in terms of  $du$ 's and  $dx$ 's).

You have just found the mean rotation angle, " $w$ ".

4) In terms of  $du$ 's and  $dx$ 's, what is (shear strain +  $w$ )? What is (shear strain -  $w$ )?

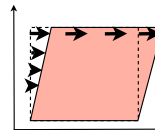
Usually we have to deal with strain that looks like this:



The  $x_1$  - parallel and  $x_2$  - parallel sides have been rotated by different amounts. You can't make this happen with a rigid transparency because rotation and strain have both occurred.

- 5a) Compute:  $du_1/dx_2$  and  $du_2/dx_1$   
5b) Compute the rotation  $w$ .  
5c) Compute the shear strain  $\epsilon_{12}$ .

6) Show me that for "simple shear" strain,  $\epsilon_{12} = w$  (or  $-w$ ).



3c) Write out the mean of the two angles in 3b (in terms of du's and dx's).

You have just found the mean rotation angle, "w".

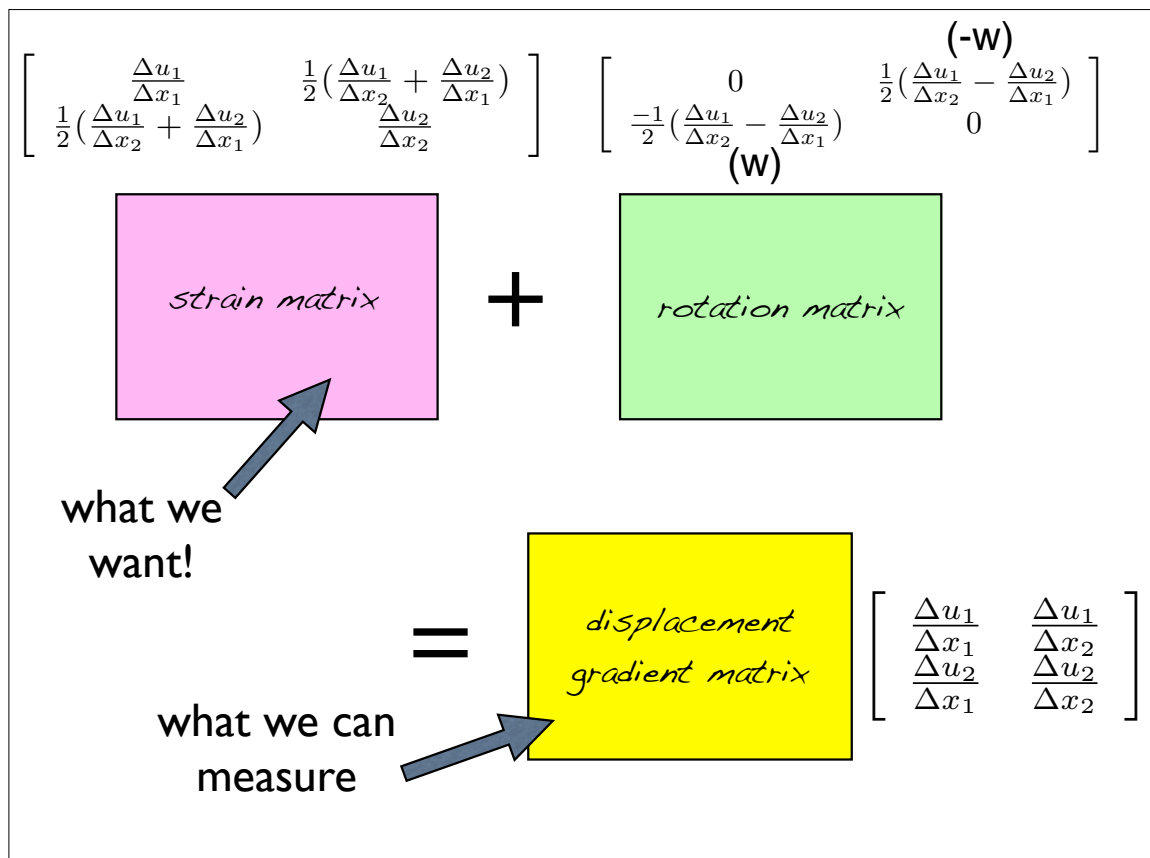
$$\frac{1}{2} \left( \frac{\Delta u_2}{\Delta x_1} - \frac{\Delta u_1}{\Delta x_2} \right)$$

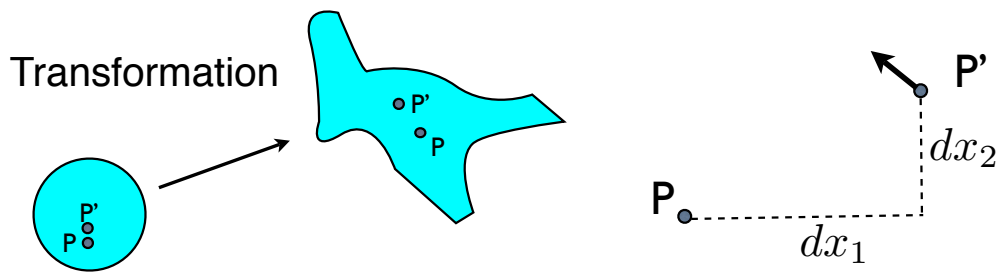
4) In terms of du's and dx's, what is (shear strain + w)?

$$\frac{1}{2} \left( \frac{\Delta u_2}{\Delta x_1} + \frac{\Delta u_1}{\Delta x_2} \right) + \frac{1}{2} \left( \frac{\Delta u_2}{\Delta x_1} - \frac{\Delta u_1}{\Delta x_2} \right) = \frac{\Delta u_2}{\Delta x_1}$$

What is (shear strain - w)?

$$\frac{\Delta u_1}{\Delta x_2}$$





$$\begin{bmatrix} du_1 \\ du_2 \end{bmatrix} = \begin{bmatrix} \text{Displacement} \\ \text{gradient} \\ \text{matrix} \end{bmatrix} \begin{bmatrix} dx_1 \\ dx_2 \end{bmatrix}$$

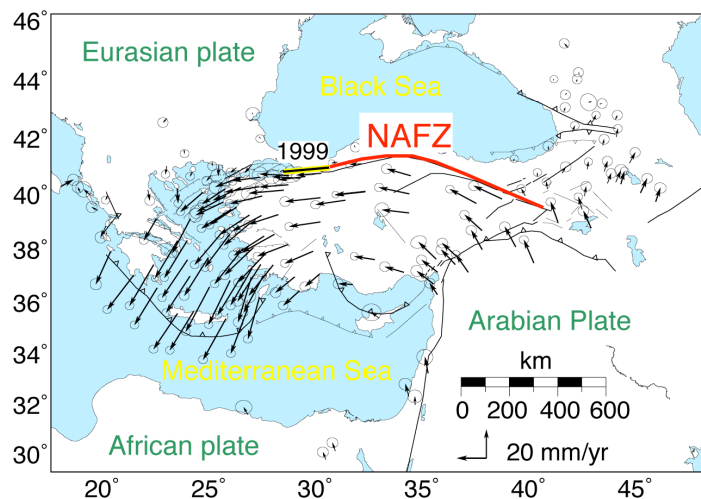
- gives displacement of P' relative to P WHICH WE CAN MEASURE

$$\begin{bmatrix} du_1 \\ du_2 \end{bmatrix}_S = \begin{bmatrix} \text{strain} \\ \text{matrix} \end{bmatrix} \begin{bmatrix} dx_1 \\ dx_2 \end{bmatrix}$$

- gives displacement of P' relative to P due to strain only (not rotation)

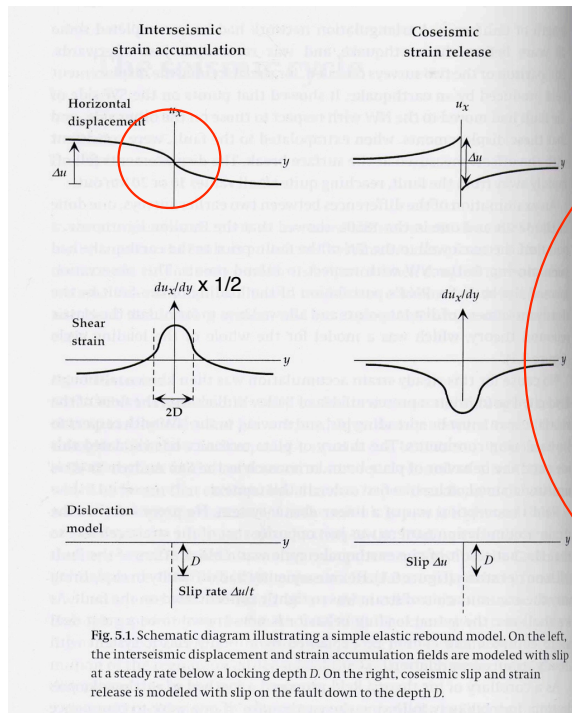
## Real-life example of **rotation**: Turkey

GPS horizontal velocities relative to Eurasia 1988-1997

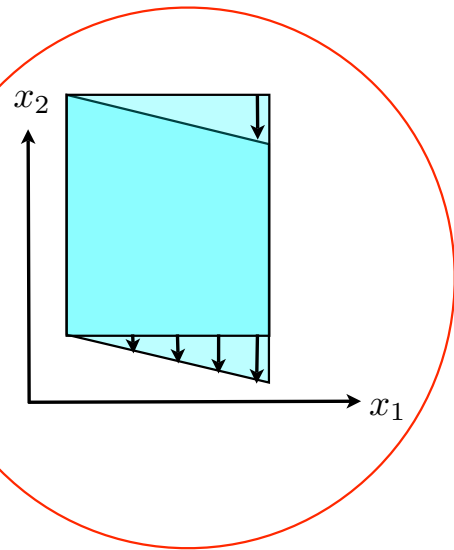


from McClusky et al., JGR, 2000

## Earthquake cycle model - strike-slip fault



From Scholz, Mechanics of Earthquakes and Faulting, 2002



Real-life example of simple shear strain around a fault

## Shear strain in one year along the San Andreas Fault



$$\frac{\Delta u_2}{\Delta x_1} \approx \left( \frac{-0.020 \text{ m}}{10000 \text{ m}} \right)$$

$$\frac{\Delta u_1}{\Delta x_1} = 0$$

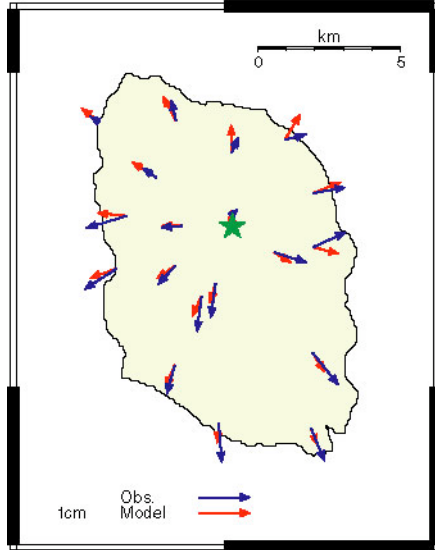
and other derivatives = 0

$$\epsilon_{12} \approx ?$$

Relative plate motion is 34 mm/year but we see only part of this relative motion close to the fault (arctan function)

Mt Fuji 2002-2003

HORIZONTAL DISPL. (wtz=0.5)  
SOURCE: H=4.3KM V=2.0E+6



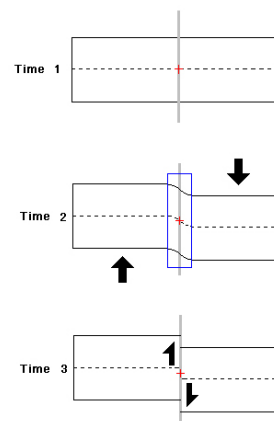
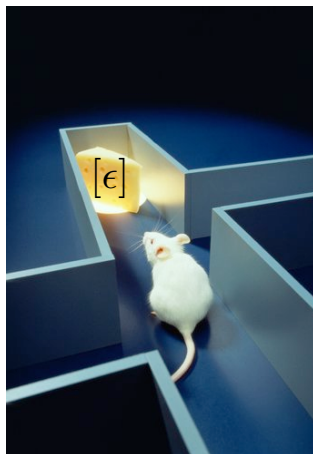
東京大学  
THE UNIVERSITY OF TOKYO  
[www.eri.u-tokyo.ac.jp](http://www.eri.u-tokyo.ac.jp)

Real-life example of  
normal strain  
around a volcano



Mt St. Helens 2004 (USGS)

For figuring out elastic force acting the fault surface,  
pushing it toward failure, **all that matters is strain**

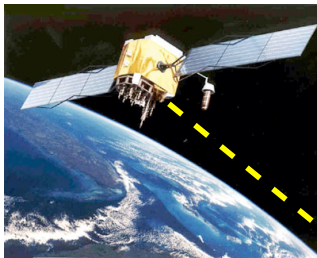
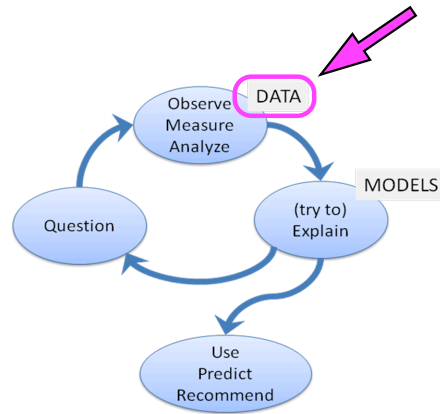


Translation and rotation do not distort the volume (*pull the spring*) and they  
**do not matter at all to us.**

GPS data give us the displacements and the displacement gradient matrix.  
**We know how to get strain from this matrix now!**

Satellite geodesy - widespread since the early 1990's:

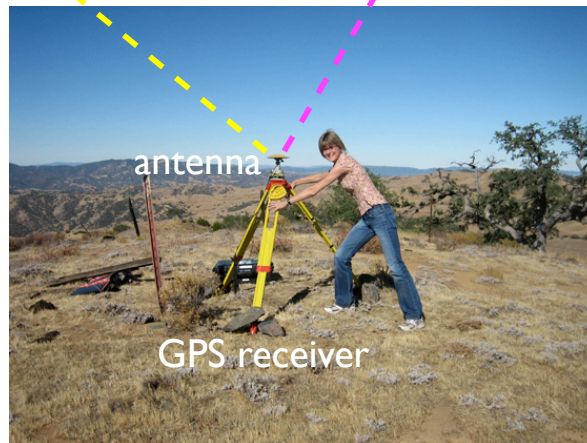
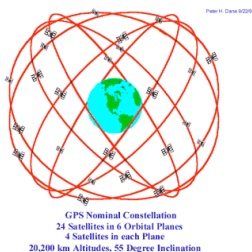
GPS  
InSAR



Satellite geodesy for  
crustal deformation  
monitoring-  
since the 1990's



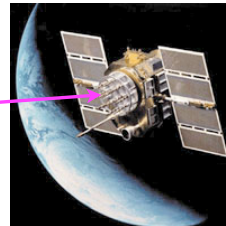
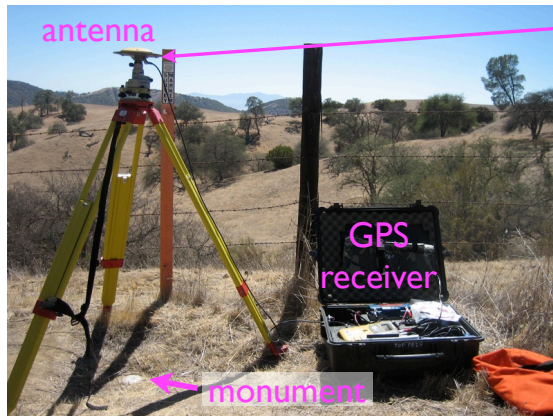
GPS satellites  
are in polar  
orbits





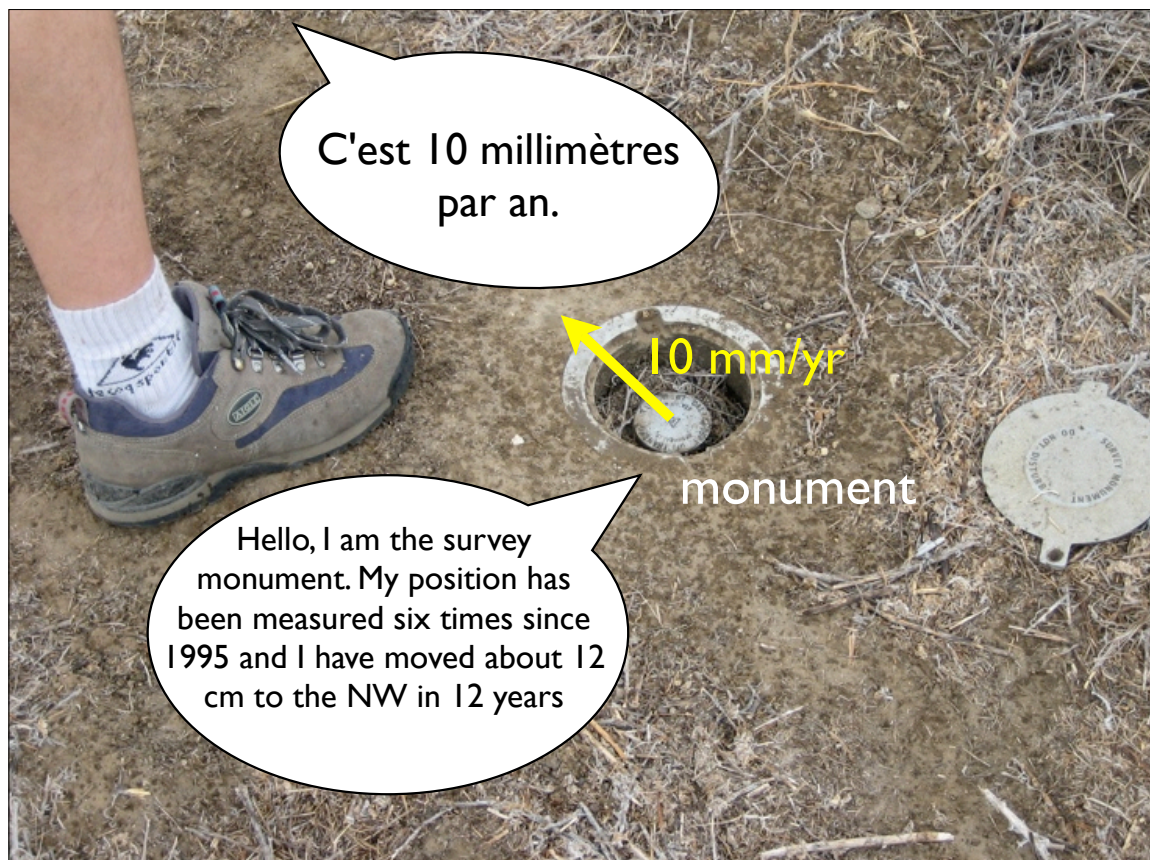
32 of these as of March 2008

(edited since lecture, not examinable - EHH)



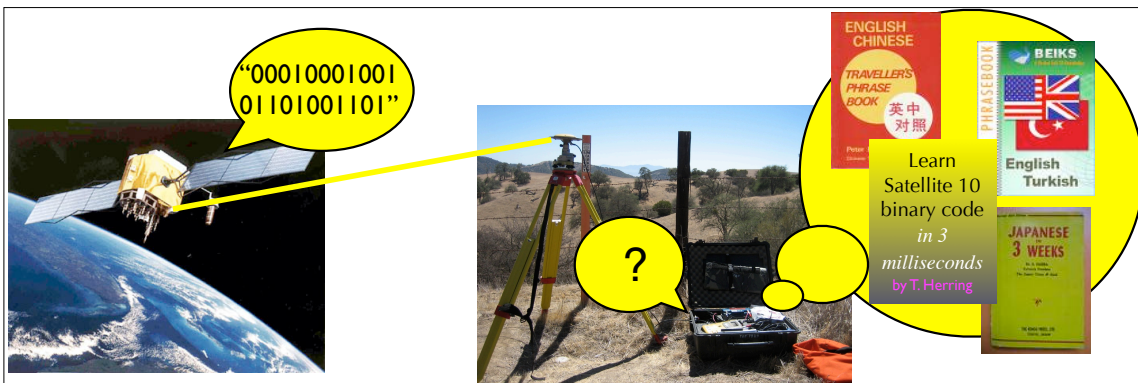
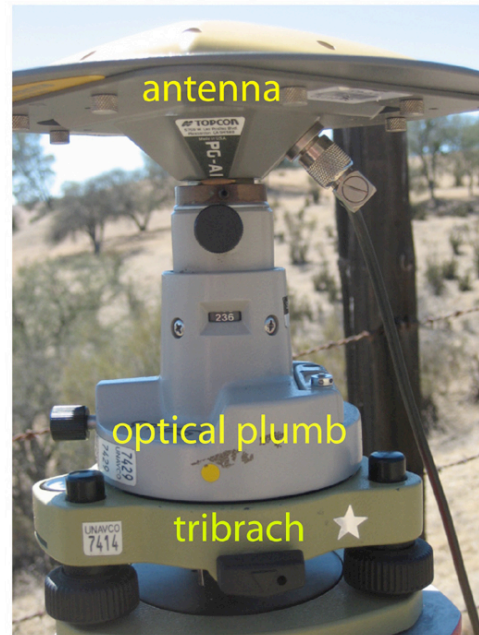
Each satellite broadcasts its own tunes at about 1500 Hz and 1200 Hz

repeated position measurements are made over time give velocity to within less than 1 mm/year!



GPS receiver - not the car kind

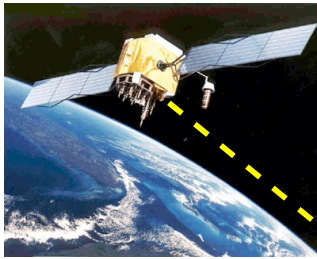
- ~ \$5,000 (with antenna etc.)
- “dual frequency” (both GPS frequencies are recorded)
- lots of memory and sophisticated software to process the data



## When GPS receiver gets a signal

- It **compares that signal with all the known codes** (there are currently 37).
- The receiver **determines which satellite it is**.
- It knows when the signal was sent (comes with a time stamp) and when it was received, multiplies the difference by the speed of light to find the distance (pseudorange).
- Once it has done that for 3 satellites, it can determine the location.





Satellite sends signal at  $t_1$ , signal is received at  $t_2$

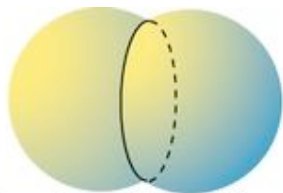
$$\text{Time Difference (in seconds)} \times 2.99792458 \times 10^8 \text{ meters/second} = \text{Distance (in meters)}$$

GPS signal tells you how far it is from the satellite ("pseudorange").

What is a typical time difference?

GPS satellites are  $\sim 2 \times 10^7$  meters above the Earth.

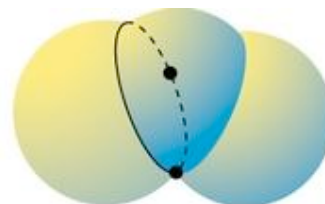
about 70 milliseconds



Two spheres intersect in a circle

Intersecting spheres of radius  $r$

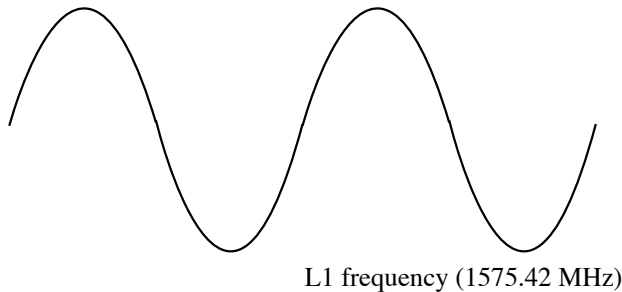
3 spheres intersect at two points - but only 1 point is on the Earth



Three spheres intersect in two points

Extra tricks are required to get position accuracy of 1 mm

Actually each satellite is broadcasting 3 tunes at once ... one really low-pitched tune with the time stamp (50 Hz) the C/A code (song) and the P(Y) code song (1 and 10 MHz)



Two microwave carrier signals.  
The L1 frequency (1575.42 MHz) carries the navigation message and the SPS code signals.  
The L2 frequency (1227.60 MHz) is used to measure the ionospheric delay by PPS equipped receivers.

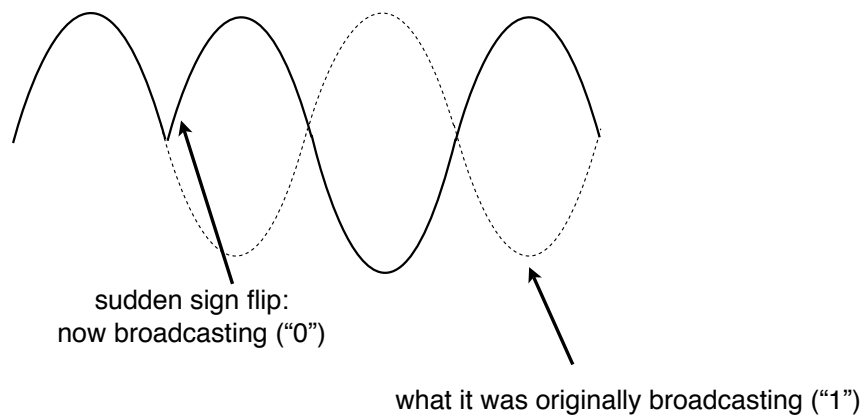
The C/A Code (Coarse Acquisition) modulates the L1 carrier phase. The C/A code is a repeating 1 MHz Pseudo Random Noise (PRN) Code. (This is the one used by your car's GPS)

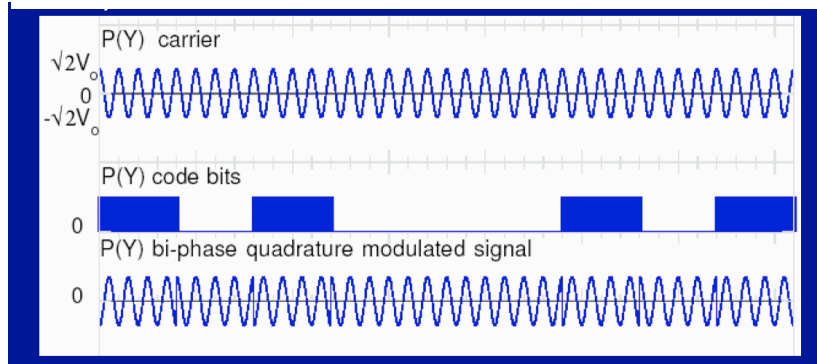
The P-Code (Precise) modulates both the L1 and L2 carrier phases. The P (Y)-Code is 10 MHz the basis for the PPS. Shorter time intervals of 0's and 1's allow a more precise pseudorange.

The only lyrics are "0" or "1" (binary code)

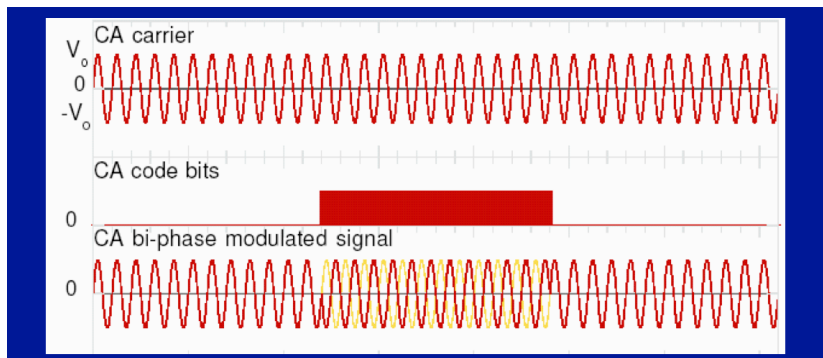
0 and 1 are indicated by a sudden flip in the polarity of the signal...

0 or 1: indicated by a sudden flip in the polarity of the signal





about 0.3  
microseconds



about 3  
microseconds

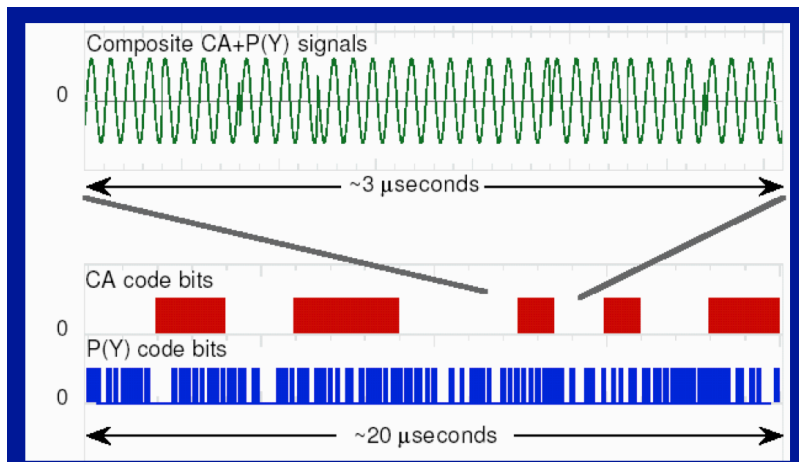
**“about 70 milliseconds” =  $2 \times 10^7$  meters. Positioning can be to within 0.001 m. How??**

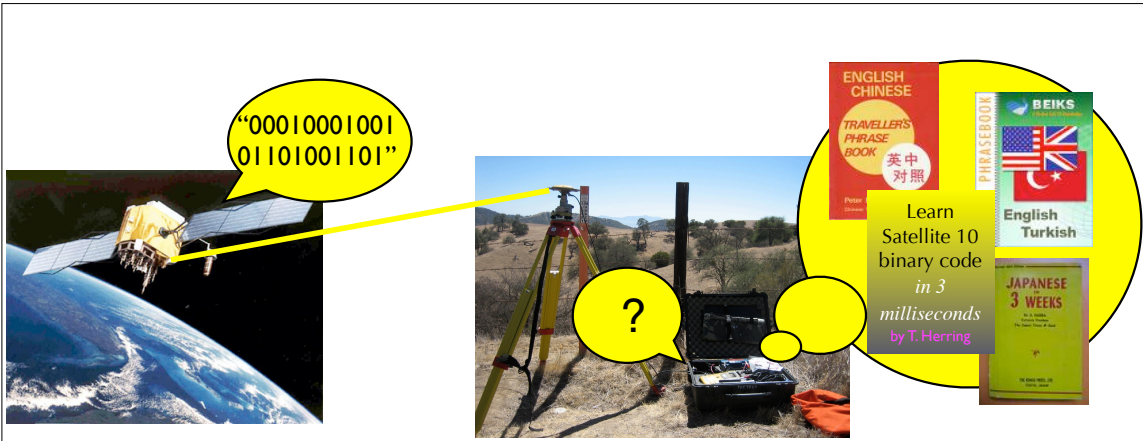
The P(Y)-code is 10 MHz the basis for precise positioning. Shorter time intervals of 0's and 1's allow a more precise pseudorange. We can also measure the phase of the signal (not the code) and get the “phase pseudorange”.

The C/A code (1 MHz) is used for coarse positioning.

In tectonic geodesy we measure position every minute or even every second for a long time, then compute a mean. This makes the measurement even more precise.

We also correct the data for weather. Water in the atmosphere slows the microwave signal down but this is corrected for using the second frequency (L2) signal.





For example, here are the first 1000 numbers of the code for satellite 1

```

00001000101001110000111001001000100001000101011000111101110010101101100111101011
00101100101001100111111011001111001001100110100011100010010001011000101101110000
00110110010001000101101000101001000000011110001100010111101111100110111001011
01111000111110101001010000101010011100001101001110110001111011110000111111111
0100100100100110011101010111100001000101010011111000010011011100111000110110
10110110101000010110100101000101001000111001110001010010111010111010101000001011
011100110110011010000000000011101110110001101010101010101000111000110011001111
01011110011101010100000111111001001010000001101000111101101001010110000010010
01001100001101100001110111011100011011101100111001000110101010000110110100101
1100101111111011000111000000110111000110000001000000110101000101011110
110001110110100011001010111100111101000000011011100110011101011110000011110110
0100010010101110000000100001010101001111011001111011001111100101111000100110101

```

## How do you compare codes?

```

100111101110100010011011111111101
-----
0000100010100111100001110010010001

```

Every time the numbers agree, add 1.  
Every time the numbers disagree, subtract 1.

This example: same satellite codes, but shifted

Not so good - score of -3.

```

01100010101011001000100100000110000011110000
11000101010110010001001000001100000111100001

```

But if you recognize they are shifted by 1:

```

01100010101011001000100100000110000011110000
11000101010110010001001000001100000111100001

```

Agreement is perfect.



Each red dot tells  
you the position of  
a GPS receiver on  
a single day.

Churchill is moving  
1.9 cm/yr west, 0.6  
cm/yr south, and 1.1  
cm/yr up.

